Investigation

Hubble’s Law

What’s the evidence that our Universe is expanding, contracting, or staying the same?

Materials

- Worksheet “Spectra of Fast-Moving Galaxies”
- Worksheet “Hubble’s Law”

The line spectrum of hydrogen, the most common element in the Universe, has characteristic lines (wavelengths or frequencies) in the regions red, turquoise, blue, and violet. The boldest line in the hydrogen spectrum is in the red region. That is if the star or galaxy is not moving towards you or away from you. If it is moving towards you or away from you, each spectrum line will be shifted either toward the red or toward the violet end of the spectrum because of the Doppler effect.

9.1. If a star is coming towards us, which end of the spectrum will its spectrum lines be shifted towards? If a star is going away from us, which end of the spectrum will its spectrum lines be shifted towards?

On the “Spectra of Fast-Moving Galaxies” data sheet there are spectra of a number of galaxies. The darkest line indicated in each the spectrum is the one that is normally in the red region of the spectrum. With extreme Doppler shifts, that bold line can appear in radically different parts of the spectrum. The scales at the top and bottom of the sheet relate Doppler shifts of the galaxies’ spectra with velocities of the galaxies. Positive velocity means the galaxy is moving away from us and negative velocity indicates the galaxy is moving towards us.

9.2. If a galaxy’s spectrum is shifted towards the red end of the spectrum, is the galaxy moving towards us or away from us?

Find out how fast each galaxy is moving. Plot the speed on the “Hubble’s Law” worksheet to create a graph of distance vs velocity.

9.3. What does the graph tell you?

The relationship between galaxy distance and velocity that you determined is called the Hubble Law because it was first discovered by astronomer Edwin Hubble.

9.4. What does Hubble’s Law imply about how our Universe is behaving?

Measuring red shift turns out to be yet another powerful way to find the distances to the most far away galaxies. Assuming Hubble’s Law applies for most galaxies, astronomers estimate distances to the most remote galaxies by measuring red shifts, finding velocities, and calculating distances from Hubble’s Law.

9.5. How distant is a galaxy that is found to be receding from us at 120,000 km/sec?
Spectra of Fast-Moving Galaxies

- **Galaxy A**: 0 (Milky Way)
- **Galaxy B**: 800 million LY
- **Galaxy C**: 400 million LY
- **Galaxy D**: 1800 million LY
- **Galaxy E**: 950 million LY
- **Galaxy F**: 2000 million LY
- **Galaxy G**: 1400 million LY
- **Galaxy H**: 700 million LY
- **Galaxy I**: 120 million LY
- **Galaxy J**: 1800 million LY
- **Galaxy K**: 1700 million LY
- **Galaxy L**: 750 million LY
- **Galaxy M**: 1400 million LY
- **Galaxy N**: 2000 million LY
- **Galaxy O**: 900 million LY
- **Galaxy P**: 1300 million LY
- **Galaxy Q**: 1700 million LY
- **Galaxy R**: 800 million LY